

WFIRST-SDT Final Report

Exoplanet Coronagraphy

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1. Introduction Kasdin, Guyon
Background on coronagraph and exoplanet science, motivate objective of showing that for minimum high-contrast case there is still excellent science and for stretch performance there is additional science, perhaps touching on SuperEarths. Also explain three main science cases: Transitional & Debris Disks, Exozodi, Giant Planets.
2. Performance assumptions Kasdin, Guyon
baseline: $3 \lambda/D$ coronagraph, $1e-8$ raw, $1e-9$ calibrated
aggressive: reducing iwa down to 2 and $1 \lambda/D$, more aggressive calibration assumptions (factors of 100 to 1000).
3. Debris Disk Science Greene
4. Exozodi measurements and effect Traub
5. Polarization science Breckenridge
6. Planets Kasdin, Greene, Traub
Radial Velocity Planet Characterization
Giant Planet discovery and Spectroscopy
Design Reference Mission Assumptions and Predictions
7. Aggressive Science/Rocky Earths & Super-Earths Guyon

AFTA WFIRST Coronagraph (WFIRST-C) Instrument Description

A Joint Report of the AFTA WFIRST Coronagraph Science Definition Team and Mission Concept Study Team

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CHANGE LOG

Version	Date	Author	Revision
0.0	11/22/05	Kasdin, Levine	First version presented at SDT #2

COMMENTS

Because of the page constraints Science: 5 pages & Design: 3 pages - details of the AFTA WFIRST Coronagraph (WFIRST-C) will be provided in appendices

The WFIRST-C study team will meet regularly (bi-weekly?) through April to monitor progress

The schedule is TBD

1/10/12 Section leads to provide rough draft

Xx/xx/xx Fill-in report milestones

4/01/12 Final draft of WFIRST-C report delivered to AFTA SDT

Writing responsibilities are defined with the following colors:

blue (SDT leads), **red** (Design team leads), **green** (together)

The goal is to describe the science objectives and requirements, the design reference mission, and the WFIRST-C baseline design concept that demonstrates feasibility of the mission. The baseline design will identify appealing science achievable with existing technology. Science achievable with a more aggressive coronagraph will be presented as stretch goals.

It is noted that WFIRST-C is “a” concept which achieves the science goals, and it is not intended to be the final design. Further studies and technology demonstrations will be required to optimize the performance and coronagraph designs. Alternative options will be described herein for reference.

The baseline concept also assumes that an order of magnitude improvement in contrast is achieved through data post-processing. Section 4 will identify techniques for improving contrast through data processing.

The document must have a logical flow and should follow the familiar pattern of science requirements -> technical requirements -> technical description -> technology development plan.

OUTLINE

1. WFIRST-C Baseline Mission Concept Description & Rationale

3.1. Introduction

Set the stage for decisions – what is important, evolution of the design & summary system description. Describe in broad terms telescope, sss, instrument ... Follow the photons. (Speckle removal technique: possible dithering, or DM options, ...), Observing scenario, timeline, orbit ...

Mission Description Table

Describe each WFIRST-C element as an existence proof: a demonstration of feasibility with the understanding that not all options have been investigated to date and trades are still work in progress.

All alternative options and trades summarized later in section 4.

FB1 validation for Dynamic EB only.

3.2. Science Derived Requirements (Shaklan Horner)

science reqts, mission reqts: sensitivity, throughput, resolution, observing overhead, orbit, observing scenario

3.3. System Architecture

3.3.0.1. Optical Configuration

3.3.0.2. Mechanical Configuration

3.3.0.3. Mechanisms

3.3.0.4. Thermal Control

3.3.0.5. Data Systems

3.3.0.6. Mass

3.3.0.7. Power

3.3.0.8. Trades

3.3.1. Science Instrument Payload

3.3.1.1. Starlight Suppression System

3.3.1.2. Wave Front Sensing and Control

3.3.1.3. Low Order Wave Front Sensor and Control

3.3.1.4. Spectrometer

3.4. Mission Operations

3.4.2. Mission Description

3.4.3. Observatory Field of Regard

3.4.4. Observational Strategy of Each Target Star

3.4.5. Orbit Environmental Issues

3.5. WFIRST-C Performance Requirements

3.5.6. Static Error Budget

3.5.7. Contrast Stability Error Budget

3.6. Baseline Observatory Performance

3.4.1. Modeling Approach (CIELO)

3.4.2. WFIRST-C Design Performance Assessment

3.4.2.1. System Thermal Performance

3.4.2.2. Contrast Performance Assessment

3.4.2.2.1. Static Performance

3.4.2.2.2. Contrast Stability Performance

3.4.3. Science Performance Assessment

Summary of what is described above in SRD and DRM, and demonstrate that WFIRST-C can accomplish the science.

3.4.3.1. Performance for exoplanet finding and characterization

3.4.3.2. Performance for giant planets and planetary system science

3.4.3.3. Performance for Circumstellar Debris Disks and Planet Formation

3.5. Verification Approach

3.5.1. tbd

3.6. Alternative Starlight Suppression Concepts and Trades

4.1.1. Data post-processing for improved contrast (Guyon)

4.1.2. PIAA-CMC (Guyon)

4.1.3. Lyot (Trauger)

4.1.4. Visible Nulling Interferometer (Shao, Clampin)

4.2. Key Technologies, TRL Assessment, Development Plan & Progress to Date

4.2.5. Tbd

4.3. Schedule

4.4. Costs

4.5. References